

## Lecture 11 : Neutrinos -1

- Role of the neutrinos in elementary particle physics
- Flavors of neutrinos : Electron, muon and tau neutrinos and their antiparticles
- The historical development of the discoveries in neutrino physics
- Solar neutrinos : Fusion chain reactions in the Sun produce neutrinos

<u>PP chain</u>	<u>Energy in MeV</u>
(pp)	
$p + p \rightarrow 2 H + e^+ + \nu_e$	$\leq 0.42$
(pep)	
$p + e^- + p \rightarrow 2 H + \nu_e$	1.552
( <sup>7</sup> Be)	
${}^7\text{Be} + e^- \rightarrow {}^7\text{Li} + \nu_e$	0.862(90%) 0.384(10%)
(hep)	
${}^2\text{He} + p \rightarrow {}^4\text{He} + e^+ + \nu_e$	$\leq 18.77$
( <sup>8</sup> B)	
${}^8\text{B} \rightarrow {}^8\text{Be}^* + e^+ + \nu_e$	$\leq 15$

<u>CNO chain</u>	<u>Energy in MeV</u>
( <sup>13</sup> N)	
${}^{13}\text{N} \rightarrow {}^{13}\text{C} + e^+ + \nu_e$	$\leq 1.199$
( <sup>15</sup> O)	
${}^{15}\text{O} \rightarrow {}^{15}\text{N} + e^+ + \nu_e$	$\leq 1.732$
( <sup>17</sup> F)	
${}^{17}\text{F} \rightarrow {}^{17}\text{O} + e^+ + \nu_e$	$\leq 1.74$

Energy spectra of solar neutrinos :

Study of the Figure 11.2 in the D.Griffiths's book. Source J.N.Bachall and A.M.Serenelli and S.Basu Astrophysical Journal, 621, L85, 2005

The great discover by J.Davis in 1968 to measure solar neutrino flux using a huge tank of chlorine in the Homestake Mine at South Dakota. Nobel Prize of the year 2002 is later awarded to J.Davis.

Two different languages :

- Flavor eigenstates of the neutrinos :  $\nu_e$  ,  $\nu_\mu$  ,  $\nu_\tau$
- Mass eigenstates of the neutrinos :  $\nu_1$  ,  $\nu_2$  ,  $\nu_3$

Mixing matrix of flavor and mass eigenstates of the neutrinos

a) Two flavor case

$$\begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix} = \begin{pmatrix} \cos \theta_{12} & -\sin \theta_{12} \\ \sin \theta_{12} & \cos \theta_{12} \end{pmatrix},$$

b) Three flavor case :  $U_{\text{PMNS}}$  mixing matrix

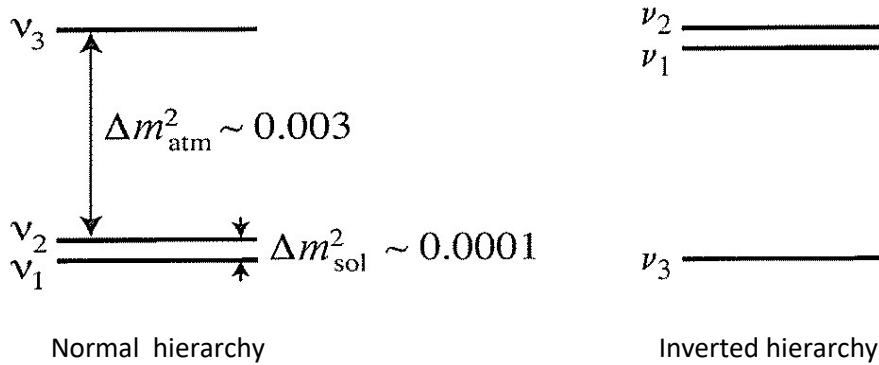
$$\begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}$$

The above  $U_{\text{PMNS}}$  mixing matrix contains three mixing angles  $\theta_{12}$  ,  $\theta_{23}$  ,  $\theta_{13}$  and CP violating Dirac phase factor  $\delta$  and Majorana phases  $\alpha_1$  and  $\alpha_2$

$$U_{\text{PMNS}} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha_1} & 0 \\ 0 & 0 & e^{i\alpha_2} \end{pmatrix}$$

The phases are not accessible by the neutrino oscillations.

The absolute values of the neutrino masses are not known; only the differences of their squares are measured via experiment. Two possible ordering of the masses are :



Transition probability between the flavors  $\nu_\alpha$  and  $\nu_\beta$  :

$$P(\nu_\alpha \rightarrow \nu_\beta) = |\langle \nu_\beta | \nu_\alpha(t) \rangle|^2 = \left| \sum_{j=1}^n \sum_{k=1}^n U_{\alpha j}^* U_{\beta k} \langle \nu_k | \nu_j(t) \rangle \right|^2$$

$$\approx \sum_{j,k} U_{\alpha j}^* U_{\beta j} U_{\alpha k} U_{\beta k}^* e^{-i\Delta m_{jk}^2 L/2E}$$

In the simple case of two flavors the above formula reduces to

$$P(\nu_\alpha \rightarrow \nu_\beta) = \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E_\nu} \right)$$

- Solve the following problems at the end of the Chapter XI of the textbook “Intr. to Elementary Particle Physics by D.Griffiths 2nd Ed. J.Wiley.”

Solve Problem 11.1

Solve Problem 11.2

Solve Problem 11.3

Solve Problem 11.4